$$\begin{aligned} & \mathcal{J}\text{-flavor mixing} \\ & \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e2} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \end{aligned}$$

$$U = \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{13}s_{23}c_{12}e^{i\delta} & c_{23}c_{12} - s_{13}s_{23}s_{12}e^{i\delta} & c_{13}s_{23} \\ s_{23}s_{12} - s_{13}c_{23}c_{12}e^{i\delta} & -s_{23}c_{12} - s_{13}c_{23}s_{12}e^{i\delta} & c_{13}c_{23} \end{pmatrix}$$

3angles $\theta_{12} \theta_{13}$ and θ_{23} and complex phase δ

 $P(\nu_{\alpha} \to \nu_{\beta}) = \left| \langle \nu_{\beta} | e^{-iH_0 L} | \nu_{\alpha} \rangle \right|^2 = \sum_{i,j} U_{\alpha i} U^*_{\beta i} U^*_{\alpha j} U_{\beta j} e^{-i\delta m^2_{ij} L/2E}$

3 masses -> only 2 mass differences If one mass difference >> other, then algebra gets simpler.

$$\Delta m_{32} \sim \Delta m_{atm}$$

 $\Delta m_{21} \sim \Delta m_{solar}$

SuperK is probably measuring:

$$P(\nu_{\mu} \to \nu_{\tau}) \simeq 4|U_{\mu3}|^{2}|U_{\tau3}|^{2}\sin^{2}(\frac{\delta m_{atm}^{2}L}{4E})$$

= $\sin^{2}(2\theta_{23})\cos^{4}(\theta_{13})\sin^{2}(\frac{\delta m_{atm}^{2}L}{4E})$

- $Sin^{2}2\theta_{23} \sim 1!$
- CHOOZ reactor measures:

 $P(\nu_e \to \nu_e) \simeq 1 - 4|U_{e3}|^2 (1 - |U_{e3}|^2) \sin^2\left(\frac{\delta m_{atm}^2 L}{4E}\right)$ $= 1 - \sin^2(2\theta_{13}) \sin^2\left(\frac{\delta m_{atm}^2 L}{4E}\right),$

• $Sin^{2}2\theta_{13} < 0.1$

Not so fast!

- LSND see a signal in v_{μ} -> v_e
- Karmen don't
- Boone 2002 will check
- 3 mass differences!
 - 1 or more experiments are wrong
 - There are more than 3 neutrino types



Sterile Neutrinos

- Know from Z width only 3 neutrinos couple to Z.
- Any additional neutrinos can mix but not show up in weak interactions
- Possible signals
 - Different MSW effects in the sun as interactions are different
 - MSW effect in atmospheric neutrinos
 - Missing vd->vd events at SNO as sterile neutrinos won't interact

SNO will tell us



Neutrino Scenario

 Sterile neutrinos will depress the neutral current (vd -> vd) rate relative to charged currents.

Prospects

Next 5-10 years

- Boone will see if there are neutrino oscillations at the $\delta m^2 \sim 0.1 \mbox{ eV}^2$ scale
- K2K will confirm (or not) SuperK, may measure δm_{23}^2
- MINOS and CERN-Gran Sasso will start high statistics measurements to $\delta m_{23}^2 \sim 0.002 \text{ eV}^2$, may see θ_{13} in v_e -> v_u
- KAMLAND will begin to explore solar neutrino land
- SNO and Borexino will see (or not see) sterile neutrinos
- We will probably know the major mass difference and mixing angles

More to do

There are 3 mixing angles and a CP violating phase

MSW effects can tell us if



If we get real lucky, we may see CP violation



 Interactions with matter can cause enhancements for v or anti-v but not both



To see CP violation, all 3 flavors must be involved. Only see it if Dm_{solar} is not too small and the Jarlskog factor J is large.

$$J = c_{12}c_{13}^2c_{23}s_{12}s_{13}s_{23}(\sin\delta)$$



Bigger v beams

- Need baselines from 2000-10,000 km to see matter and CP effects
- Super MINOS?
 - Conventional beam with higher energy, more flux
- Neutrino Factory
 - Muon storage ring $\mu \rightarrow e \nu_{\mu}$ anti- ν_{e}

Bigger v detectors

- NNN workshops
 - Next one at FNAL Aug 7-8
- Consider a neutrino detector 10 times bigger than Super K
 - 1000 neutrinos from a supernova
 - Proton decay
 - Neutrino astrophysics
 - Solar neutrinos
 - Target for long baseline accelerators

Conclusions

- 5 years ago neutrino oscillations were a possibility
 - More solar neutrino data at different energies
 - Better solar models
 - Super K atmospheric neutrinos
 - Confirmation from MACRO, SoudanII ...
- They are now an industry
 - K2K
 - MINOS
 - CGS
 - KamLand
 - SuperDuperK
 - Neutrino factories?